## Atmospheric transport and dispersion modelling study of the I-131 detected in Jan/Feb 2017 in Europe

Pieter De Meutter<sup>1,2</sup>, Johan Camps<sup>1</sup>, <u>Andy Delcloo<sup>2,3</sup></u> and Piet Termonia<sup>2,3</sup>

Pieter.Demeutter@sckcen.be

<sup>1</sup>: Belgian Nuclear Research Institute

<sup>2</sup>: Royal Meteorological Institute of Belgium

<sup>3</sup>: Ghent University, Department Physics and Astronomy

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- 1. Inverse atmospheric transport modelling: first attempts
- Direct atmospheric transport modelling: using release assumptions, can we reconstruct the <sup>131</sup>I detections?
- 3. Effect of the meteorological conditions
- 4. Inverse atmospheric transport modelling revisited
- 5. Conclusions

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#### In January and February 2017, <sup>131</sup> was detected throughout Europe; its origin is not fully understood



Masson et al., (2018). Potential Source Apportionment and Meteorological Conditions Involved in Airborne 131I Detections in January/February 2017 in Europe. Environmental Science & Technology, 52(15), 8488-8500.

## Inverse atmospheric transport modelling: a three-step problem

#### 1. Input data

Numerical weather prediction data:

ECMWF IFS: 3-hourly data coarse-grained to 1° horizontal grid spacings

#### Iodine-131 observations:

28 detections from the Ro5 + detections from the CTBTO IMS radionuclide network

#### 2. Atmospheric transport and dispersion modelling

The Lagrangian particle model Flexpart in backward mode *(Seibert and Frank,* 2004)

## Source-receptor relationship:

Flexpart calculates the source-receptor-sensitivities  $M_{ij}$  for each observation  $y_i$ :  $y_i = M_{ij}x_j$ 

#### 3. Inverse modelling

A source term  $x_j$  is found by minimizing a cost function:

$$exp\left(\frac{1}{n}\sum_{i}\left(\log(y_i+\alpha)-\log(M_{ij}x_j+\alpha)\right)^2\right)$$

The optimisation is solved using a quasi-Newton technique and does not require to rerun Flexpart.

The inverse modelling is applied to each grid box separately (single grid box source).

# Source localization based on all available observations (single source assumption)



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#### Potential sources of the <sup>131</sup>I detections



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#### Release assumptions:

Polatom, Mallinckrodt, UPRA following Table 1 of Masson et al., 2018 IRE: 1 GBq/y (FANC report) IoI: 100 GBq/y (initial assumption)

Karpov institute: 150 GBq/y (inverse modelling using CTBTO observations):

- Case 1: three detections at RN61 (12, 14, 18 January), one detection at RN54 (13 January)
- Case 2: six detections at RN61 (30, 31 January; 1, 4, 5, 6 February)
- Case 3: four detections at RN61 (17, 24, 25 February; 1 March)



A proxy for local sources: area source proportional to the population density, totaling 30 GBq/y (population density data: 1° resolution from the NASA EOSDIS database; release amount based on Fig. 2 of Masson et al., 2018)



#### <sup>131</sup>I activity concentration time series



Very poor agreement; possible reasons:

- Errors in the meteorological data
- Errors in the atmospheric transport and dispersion processes
- Errors in the emission assumptions

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## Assessing the effect of meteorology directly

- The episode of <sup>131</sup>I detections was associated with strong temperature inversions that deteriorate mixing in the lower troposphere
- In general, temperature inversions were present in the meteorological data; however, the strength of the inversion was generally underestimated (see Figure)
- Initial study suggest that in Flexpart, the parametrization of atmospheric transport and dispersion processes mainly depends on the height of the planetary boundary layer; an inversion does not seem to play a direct role



Essen

## Assessing the effect of meteorology indirectly

- Flexpart forward simulation during one full year 2017
- Release: area source with constant emission proportional to the density population (as proxy for local sources)
- Output:
  - 3-hourly simulated concentrations are converted into weekly simulated concentration (by averaging and applying a decay correction)
  - 51 simulated concentrations per station in 2017
- A quantile plot is made for each station; the simulated concentration is marked by '+' when a detection took place



Variation in concentration only depends on the meteorological conditions

## Certain detections took place when Flexpart predicts a maximum influence from local sources



Results suggest that certain detections can be explained by exceptional meteorological conditions rather than unusual emissions

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Source localization: subset of 9 observations that can be explained by exceptional meteorological conditions



#### Source localization: subset of 11 observations



#### Source localization: subset of 7 observations



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#### Conclusions

- In January and February 2017, <sup>131</sup>I was detected throughout Europe; its origin is not fully understood
- Determination of the origin is hard because there are many possible sources with time-varying emissions → many more degrees of freedom than observations
- Comparing simulated activity concentrations obtained from direct modelling with observed activity concentration leads to a poor agreement: likely we have insufficient knowledge of the emissions, in particular related to peak releases from local sources
- Results suggest that part of the detections can be linked to the exceptional meteorological conditions
- Inverse modelling results suggest that part of the detections can be linked to releases from Polatom and Karpov